Complete classification of the FLRW solutions with linear EOS: p.p. curvature singularities for general geodesics

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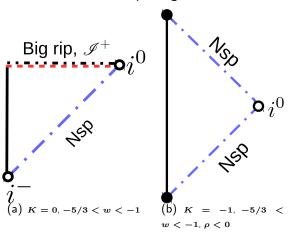
GR23 @ Beijing (hybrid), 03-08/07/2022 Harada, Carr, Igata, CQG 35 (2018) 105011 [arXiv:1801.01966] Harada, Igata, Sato, Carr, to appear in CQG [arXiv:2110.13421].

Introduction

- FLRW spacetimes
 - Spatial curvature $K=0,\pm 1$
 - EOS p=w
 ho can be assumed.
- Curvature singularities
 - Singularities may be defined by incomplete geodesics.
 - P.p. singularities: At least one component of the Riemann tensor in the parallely propagated frame is unbounded along the incomplete geodesic. The spacetime cannot be extended.
 - S.p. singularities: At least one scalar polynomial of the Riemann tensor is unbounded along the incomplete geodesic.
 Big-bang singularities are included.
- We classify FLRWs with $p = w\rho$ without assuming any energy conditions. w = -5/3 is an important critical value.

N.s.p. singularities

- N.s.p. singularities: P.p. singularities which are not s.p. Both ρ and p are bounded.
- (a) Initial null n.s.p. singularity
- (b) Initial and final null n.s.p. singularities



Big-rip singularities

- Big-rip singularities: s.p. singularities with $ho o \infty$ and the scale factor $a o \infty$ at $t o t_s$
- (c) Big-rip singularity for timelike geodesics but \mathscr{I}^+ for null geodesics
- (d) Big-rip singularity for both timelike and null geodesics
- (e) Both future and past big-rip singularities

